Molecular Dynamics and Rovibrational Spectra in a Magnetic Field

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Dynamics and Rovibrational Spectra in a Magnetic Field

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Introduction

- We have over the last decade studied atoms and molecules in (ultra)strong magnetic fields
 - ightharpoonup one atomic unit of magnetic field strength B corresponds to

$$B_0 = 2.35 \times 10^5 \,\mathrm{T}$$

- Such field strengths have not been realized on Earth (yet) but exist on magnetic white dwarf stars
 - atoms have been observed under such conditions but not vet molecules
 - Matter in Strong Magnetic Fields', D. Lai, Rev. Mod. Phys. 73, 629 (2001)
- ▶ Under such conditions chemistry becomes very different it is an exotic, largely unexplored world
 - it gives a fresh perspective on Earth chemistry and new insight into molecular electronic structure
 - it describes atoms (and molecules) observed in astrophysics
 - it is a stress test for quantum chemistry in a different environment
- ▶ We here consider molecules in ultrastrong magnetic fields with emphasis on molecular dynamics
 - background
 - molecular bonding and structure
 - molecular dynamics with the (screened) Lorentz force
 - rovibrational spectra
 - harmonic approximation

Three field regimes

▶ The electronic Hamiltonian contains a field-free term H₀ and terms that are linear and quadratic in the field

$$H = H_0 + \frac{1}{2}BL_z + Bs_z + \frac{1}{8}B^2(x^2 + y^2)$$

- lt is then useful to distinguish three field regimes:
 - Coulomb regime: $B \ll B_0$
 - Coulomb interactions dominate
 - ► familiar chemistry of spherical atoms
 - ► Earth: $10^{-10}B_0$, NMR $10^{-4}B_0$, pulsed fields $10^{-2}B_0$
 - 2 Intermediate regime: $B \approx B_0$
 - Coulomb and magnetic interactions compete
 - exotic chemistry of small, ellipsoidal atoms
 - ▶ magnetic white dwarfs: 0.01–1.0*B*₀
 - 3 Landau regime: $B \gg B_0$
 - magnetic interactions dominate
 - alien chemistry of tiny, needle-like atoms
 - neutron stars: $10^3 10^4 B_0$





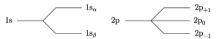


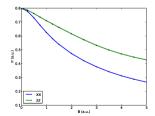
Molecular Hamiltonian – para- and diamagnetism

▶ The Hamiltonian contains linear paramagnetic terms and a quadratic diamagnetic term:

$$H = H_0 + \frac{1}{2}BL_z + Bs_z + \frac{1}{8}B^2(x^2 + y^2)$$
paramagnetic Zeeman terms diamagnetic term

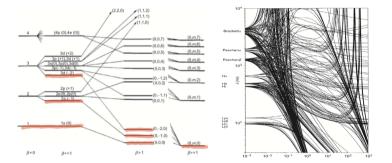
- Paramagnetic Zeeman operators:
 - couple the field to the angular momentum of the electrons
 - reduce symmetry and split energy levels
 - raise or lower the energy, depending on orientation
- Diamagnetic operator:
 - arises from induced angular momentum
 - raises the energy of all systems
 - squeezes all systems
 - ▶ helium transversal size $\propto 1/\sqrt{B}$
 - ▶ helium longitudinal size $\propto 1/\log B$
- All field operators favour high angular momentum





From Coulomb to Landau regime

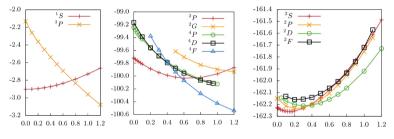
- ▶ The splitting of levels observed in the laboratory is only the thin end of the wedge
- \triangleright As the field increases, a complete state reordering occurs, driven by the magnetic quantum number m_{ℓ}
 - left: hydrogen energy levels in units of $\beta = B/B_0$ (Simola and Virtamo, 1978); right: wavelength spectrum against $\beta/2$ (Wunner and Ruder, 1987)



- The Landau levels are quantized as a harmonic oscillator except in the field direction
 - the lowest $m_{\ell} < 0$ component of each hydrogen level enters the infinitely degenerate Landau ground state

Evolution of atomic states in a magnetic field

- Lowest states of helium (left), fluorine (middle), and sodium (right) in a magnetic field
 - Stopkowicz, Gauss, Lange, Tellgren, and Helgaker, JCP 143, 074110 (2015)

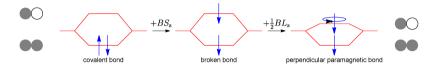


$$\langle H \rangle = \langle H_0 \rangle + \frac{B}{8} \langle S_z + \frac{1}{2} L_z \rangle + \frac{1}{8} \frac{B^2}{8} \langle x^2 + y^2 \rangle$$

- Convex evolution of energy levels with increasing field strength:
 - the initial slope is determined by the permanent magnetic moment through the Zeeman term
 - the slope increases monotonically with increasing field strength by the quadratic diamagnetic term
 - all curves therefore have a convex shape with a unique minimum, at a characteristic field strength (here lowest for sodium)
 - beyond this minimum, the energy increases indefinitely first quadratically, then linearly

Zeeman interaction: bond breaker and bond maker

► Let us consider the ground-state H₂ molecule in an increasing magnetic field



- The spin Zeeman interaction Bs_z breaks the covalent bond
 - the magnetic field favours beta spin over alpha spin
 - reventually, the alpha electron undergoes a spin flip, forcing it into the antibonding orbital (singlet-triplet transition)
 - we expect the resulting molecule to be unstable and fall apart . . .
- ▶ The orbital Zeeman interaction $\frac{1}{2}BL_z$ makes a paramagnetic bond
 - however, $1\sigma_{ii}^*$ has the shape of a $2p_{-1}$ orbital, if properly oriented
 - the $1\sigma_{ii}^*$ energy is increasingly lowered as the atoms come together
 - by this mechanism, the triplet $1\sigma_g 1\sigma_g^*$ configuration becomes bonding

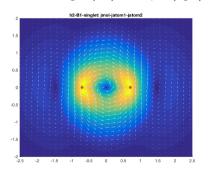


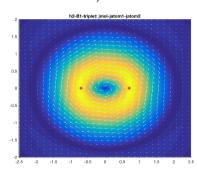


- Perpendicular paramagnetic bonding
 - Lange, Tellgren, Hoffmann and Helgaker, Science 337 327 (2012)

Diamagnetic and paramagnetic currents

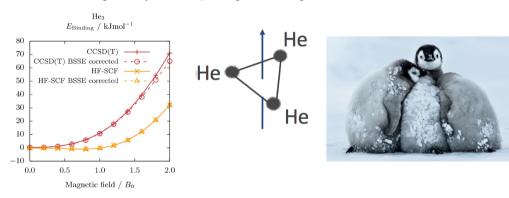
- ► The field sets up counter-clockwise currents in each atom
 - these diamagnetic currents increase the energy of each atom
- As the atoms come closer together, molecular currents appear
 - in singlet H₂, these are counter clockwise, raising the energy
 - in triplet H₂, these are clockwise, lowering the energy
- Induced currents in singlet (left) and triplet (right) H₂ (relative to free atoms)





Bonding in helium trimer

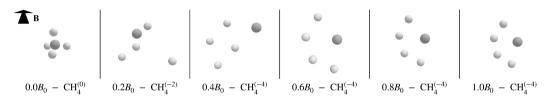
Electron correlation significantly enhances paramagnetic bonding



- Stopkowicz, Gauss, Lange, Tellgren, and Helgaker, JCP 143, 074110 (2015)
- ▶ Bond creation: clustering atoms shield themselves from the field in the same way that huddling penguins protect themselves against the cold in the Antarctic winter.

Methane from tetrahedron to fan shape

- Parmagnetically bound molecules have structures vastly different from those on Earth
- CH₄, for example, is no longer tetrahedral but fan-shaped (half moon) in a sufficiently strong field



► The CDFT dissociation energy of lowest dissociation pathway

B/B_0	0.0	0.2	0.4	0.6	0.8	1.0
$\Delta E/kJ\;mol^{-1}$	546.7	0.3	5.2	32.0	41.9	49.3
mean C–H distance (Å)				2.46	2.14	1.95
mean H–H distance (Å)				1.87	1.59	1.41

Pemberton, Irons, Helgaker, and Teale, J. Chem. Phys. 156, 204113 (2022)

Molecular dynamics in a magnetic field

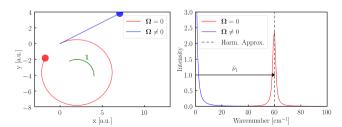
- ▶ To calculate spectra in a field, we have carried out semi-classical molecular dynamics calculations
 - it is then necessary to integrate the classical equations motion with the Lorentz force acting on each atom

$$\mathbf{F}(\mathbf{r}, \mathbf{v}) = z\mathbf{E}(\mathbf{r}) + z\mathbf{v} \times \mathbf{B}(\mathbf{r}) \quad \leftarrow \text{velocity-dependent Lorentz force}$$

▶ Each nucleus experiences a local field, which is the total field screened by the electrons:

$$\mathbf{B}_{\mathsf{loc}}(\mathbf{r}) = \mathbf{B} + \mathbf{B}_{\mathsf{el}}(\mathbf{r}) \quad \leftarrow \mathsf{screened} \ \mathsf{magnetic} \ \mathsf{field}$$

- ▶ Illustration: calculations on the He atom with and without screening
 - without screening: ${\bf B}_{\rm loc}({\bf r})={\bf B}$, giving circular motion with frequency $\tilde{\nu}_1\approx 60\,{\rm cm}^{-1}$ at $B=B_0$
 - with screening: $B_{loc}(\mathbf{r}) = 0$, giving perfect translation



Born-Oppenheimer approximation in a magnetic field

► Molecular Hamiltonian in a magnetic vector potential A:

$$H_{\mathsf{mol}} = rac{1}{2} \sum_I M_I^{-1} \, \Pi_I^2 + H_{\mathsf{el}}(\mathbf{r},\mathbf{R},\mathbf{A}), \quad \mathbf{\Pi}_I = \mathbf{P}_I - Z_I e \mathbf{A}(\mathbf{R}_I), \quad \mathbf{P}_I = -\mathrm{i}\hbar oldsymbol{
abla}_I$$

- $lackbox{lack}$ electronic coordinates ${f r}=\{{f r}_1,{f r}_2,\dots\}$ and nuclear coordinates ${f R}=\{{f R}_1,{f R}_2,\dots\}$
- Born-Oppenheimer wave-function ansatz:

$$\Psi(\mathbf{r},\mathbf{R},t) = \Theta(\mathbf{R},t)\Phi(\mathbf{r};\mathbf{R}) \quad \leftarrow \text{ nuclear wave function} \times \text{electronic wave function}$$

Project molecular Schrödinger equation against the electronic wave function to obtain nuclear Schrödinger equation

$$i\hbar \partial_t \Psi(\mathbf{r}, \mathbf{R}, t) = H_{\text{mol}} \Psi(\mathbf{r}, \mathbf{R}, t) \implies i\hbar \partial_t \Theta(\mathbf{R}, t) = (T + U)\Theta(\mathbf{R}, t)$$

lacktriangle The nuclear Hamiltonian T+U has contributions from nonadiabatic coupling matrix elements (NACMEs):

Born-Oppenheimer dynamics with the Lorentz force

The nuclear Schrödinger equation for a molecule in a magnetic field is given by:

$$\frac{1}{2} \sum\nolimits_I M_I^{-1} \left(\mathbf{P}_I - Z_I e \mathbf{A}_I + \langle \mathbf{\Phi} | \mathbf{P}_I | \mathbf{\Phi} \rangle \right)^2 \Theta(\mathbf{R}, t) + U(\mathbf{R}) \Theta(\mathbf{R}, t) = \mathrm{i} \hbar \partial_t \Theta(\mathbf{R}, t)$$

ightharpoonup The corresponding classical equations of motion for nucleus I at \mathbf{R}_I become:

$$M_I\ddot{\mathbf{R}}_I = -\underbrace{\boldsymbol{\nabla}_I U(\mathbf{R})}_{\text{gradient force}} - \underbrace{Z_I e\,\mathbf{B} \times \dot{\mathbf{R}}_I}_{\text{Lorentz force}} + \underbrace{\sum_J \boldsymbol{\Omega}_{IJ}(\mathbf{R})\,\dot{\mathbf{R}}_J}_{\text{Berry force}} \quad \leftarrow \text{screened Lorentz force}$$

The Berry force provides the screening and is obtained from the Berry-curvature tensor:

$$\Omega_{IJ}(\mathbf{R}) = 2\hbar \operatorname{Im} \langle \nabla_I \Phi | \nabla_J^{\mathsf{T}} \Phi \rangle \leftarrow \mathsf{Berry\ curvature}$$

- evaluation by response theory, making trajectory calculations in a field more expensive
- LONDON is the first general code for molecular dynamics in a magnetic field
 - Tanner Culpitt, Laurens Peters, Erik Tellgren, and Trygve Helgaker, J. Chem. Phys. 155, 024104 (2021)
 - Laurens Peters, Tanner Culpitt, Laurenz Monzel, and Trygve Helgaker, J. Chem. Phys. 155, 024105 (2021)
 - Ceresoli, Marchetti, and Tosatti, Phys. Rev. B 75, 161101 (2007) perpendicular orientation, minimal Slater basis

Berry curvature and magnetic field

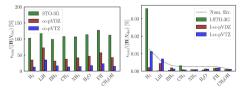
- ▶ The Berry curvature contains complete information about the magnetic field
- if not, neutral atoms would undergo cyclotron motion in a magnetic field
- ▶ The important relation is the magnetic-translational sum rule:

$$2\hbar \operatorname{Im} \sum_{IJ} \langle \boldsymbol{\nabla}_I \boldsymbol{\Phi} | \boldsymbol{\nabla}_J^\mathsf{T} \boldsymbol{\Phi} \rangle = e N_{\mathsf{el}} \begin{pmatrix} 0 & -B_z & B_y \\ B_z & 0 & -B_x \\ -B_y & B_x & 0 \end{pmatrix} \quad \leftarrow \mathsf{wave function \ determines \ the \ field}$$

- Peters, Culpitt, Tellgren, and Helgaker, JCP 157, 134108 (2022)
- ► The sum rule is a consequence of the translational symmetry of the electronic wave function

$$\Phi(\mathbf{r}; \mathbf{R} + \mathbf{T}) = \exp\left[-\frac{\mathrm{i}e}{2\hbar}\mathbf{B} \times \mathbf{T} \cdot (\mathbf{r} - \mathbf{R})\right] \Phi(\mathbf{r} - \mathbf{T}; \mathbf{R}) \leftarrow \text{translation of nuclear coordinates by } \mathbf{T}$$

▶ It holds in a finite basis only if London atomic orbitals (GIAOs) are used:

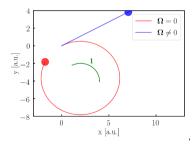


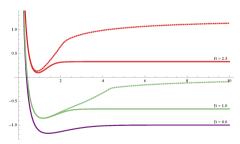
London orbitals

London atomic orbitals are Gaussians with built-in field oscillations

$$\omega(\mathbf{r};\mathbf{R}) = \exp\left[-\frac{\mathrm{i}e}{2\hbar}\mathbf{B}\times(\mathbf{R}-\mathbf{O})\cdot\mathbf{r}\right]G(\mathbf{r}-\mathbf{R})$$
 \leftarrow London atomic orbital at \mathbf{R} with gauge origin \mathbf{O}

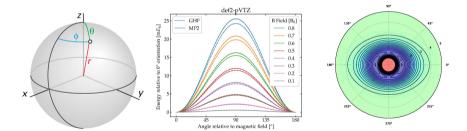
- they ensure gauge-origin invariance of all calculations
- Use of London orbitals is essential to ensure proper chemistry in a magnetic field for example,
 - translation of neutral atom in a uniform magnetic field
 - correct dissociation of molecules in a magnetic field





Rovibrational coordinates for singlet H₂

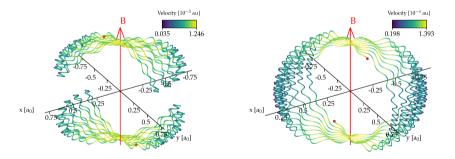
- Singlet H₂ has a preferred parallel field orientation, with a field-dependent barrier at perpendicular field orientation
- There are three degrees of freedom, in addition to translation:
 - ϕ : azimuthal motion: rotation about field axis
 - \bullet : polar motion: hindered rotation or libration (wagging)
 - r: radial motion: bond stretching



For singlet H₂, the barrier is equatorial, occurring in the xy plane

Hindered rotation vs. libration of H₂

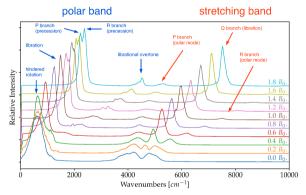
- A molecule may or may not have sufficient energy to cross the equatorial barrier, giving very different dynamics
 - trajectories initiated at equilibrium with energies 2.8 mH give libration (left)
 - trajectories initiated at equilibrium with energies 3.5 mH give hindered rotation (right)



- Lorentz force acting on the partially screened nuclei induces precession of the atoms about the field axis
- Monzel, Pausch, Peters, Tellgren, Helgaker, and Klopper, J. Chem. Phys. 157, 054106 (2022)

Rovibrational spectrum of H₂ at 1500 K

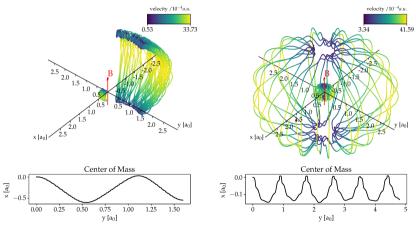
- NVT ensemble averaged over 2000 trajectories of 1 ps duration with initial velocities from Maxwell-Boltzmann distribution
 - power spectrum by Fourier transformation of velocities recorded during the simulation intensities reflect how often a mode is visited during the simulation



- polar band:
 - lacktriangle initial (red-shifted) hindered rotation followed by (blue-shifted) libration at about 0.4 B_0 overtone is visible
- coupling to precession gives P and Q branches in a strong field
- stretching band:
 - lacktriangle coupling to rotation gives P and R branches at zero field Q branch arises by coupling to libration from about $0.4B_0$
 - blue shift occurs as the bond becomes stronger with increasing field strength increased P–R splitting reflects the blue shift of libration

Trajectories of polar LiH in a magnetic field

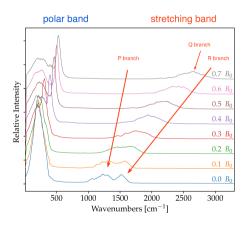
lacktriangle Trajectories with libration (left) and hindered rotation (right) of LiH at $0.2B_0$



LiH has a preferred perpendicular orientation and avoids the parallel orientation altogther
LiH has a permanent dipole moment, which induces centre-of-mass motion

Rovibrational spectrum of LiH at 1500 K

▶ NVT ensemble averaged over 2000 spectra of 1 ps duration with initial velocities from Maxwell–Boltzmann distribution



Laurens Peters, Tanner Culpitt, Laurenz Monzel, Erik Tellgren, and Trygve Helgaker, J. Chem. Phys. 155, 024105 (2021) Laurenz Monzel, Ansgar Pausch, Laurens Peters, Erik Tellgren, Trygve Helgaker, and Wim Klopper, J. Chem. Phys. 157, 054106 (2022)

Harmonic approximation and normal modes

 \triangleright Field-free harmonic rovibrational frequencies and normal modes are obtained from an $3N \times 3N$ eigenvalue problem

$$(\mathbf{H} - \omega^2 \mathbf{M}) \, \boldsymbol{\eta}_{\omega} = \mathbf{0}$$

- ► H and M are the real symmetric Hessian (force-constant) and nuclear-mass matrices, respectively
- ightharpoonup a quadratic eigenvalue problem where $(\pm\omega,\eta_\omega)$ describe excitations and deexcitations
- In a magnetic field, the quadratic eigenvalue problem contains both linear and quadratic terms:

$$(\mathbf{H} - \mathrm{i}\omega\mathbf{\Lambda} - \omega^2\mathbf{M})\,\boldsymbol{\eta}_{\omega} = \mathbf{0}, \quad \Lambda_{I\alpha,J\beta}(\mathbf{R}) = Z_I e\,\delta_{IJ}\epsilon_{\alpha\beta\gamma}B_{\gamma}(\mathbf{R}) + \Omega_{I\alpha,J\beta}(\mathbf{R})$$

- ▶ the real antisymmetric screened Lorentz matrix Λ accounts for the screened Lorentz force
- lacktriangle the 6N solutions occur in complex conjugate pairs describing excitations and deexcitations

$$(\omega, \boldsymbol{\eta}_{\omega}) \leftrightarrow (-\omega^*, \boldsymbol{\eta}_{\omega}^*)$$

The quadratic eigenvalue problem can be linearized in different ways by doubling the dimensions, for example:

$$\begin{pmatrix} \mathbf{0} & \mathbf{I} \\ -\mathbf{H} & \mathbf{\Lambda} \end{pmatrix} \begin{pmatrix} \boldsymbol{\eta}_{\omega} \\ \boldsymbol{\gamma}_{\omega} \end{pmatrix} = \mathrm{i}\omega \begin{pmatrix} \mathbf{I} & \mathbf{0} \\ 0 & \mathbf{M} \end{pmatrix} \begin{pmatrix} \boldsymbol{\eta}_{\omega} \\ \boldsymbol{\gamma}_{\omega} \end{pmatrix} \quad \leftarrow \quad 6N \times 6N \text{ linear eigenvalue problem}$$

► Tellgren, Culpitt, Peters, and Helgaker, J. Chem. Phys. 158, 124124 (2023)

Distinct eigenvalues belonging to the same eigenvector

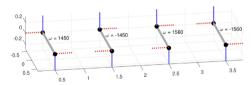
In the quadratic eigenvalue problem, distinct eigenvalues $\omega_1 \neq \omega_2$ may have the same eigenvector η :

$$\left(\mathbf{H} - \mathrm{i}\omega_1\mathbf{\Lambda} - \omega_1^2\mathbf{M}\right)\mathbf{\eta} = \mathbf{0}, \quad \left(\mathbf{H} - \mathrm{i}\omega_2\mathbf{\Lambda} - \omega_2^2\mathbf{M}\right)\mathbf{\eta} = \mathbf{0},$$

This happens if and only if η is a simultaneous eigenvector of **H** and i Λ :

$$(\mathbf{H} + \omega_1 \omega_2 \mathbf{M}) \, \boldsymbol{\eta} = \mathbf{0}, \quad (i\boldsymbol{\Lambda} + (\omega_1 + \omega_2) \mathbf{M}) \, \boldsymbol{\eta} = \mathbf{0}$$

▶ Illustration: hindered rotation of singlet H_2 at $B = B_0$



- ▶ the frequencies 1450 and -1560 cm⁻¹ belong to η , while -1450 and 1560 cm⁻¹ belong to η^*
- \triangleright at the same time η satisfies two linear eigenvalue problems

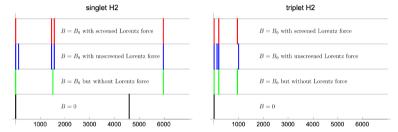
$$\mathbf{H}\boldsymbol{\eta} = (1504\,\mathrm{cm}^{-1})^2\mathbf{M}\boldsymbol{\eta} \leftarrow \text{degenerate frequency in the limit of complete screening}$$

$$i\mathbf{\Lambda}\boldsymbol{\eta} = 110 \, \mathrm{cm}^{-1} \mathbf{M}\boldsymbol{\eta}$$

← splitting of degenerate frequencey by the screened Lorentz force

Harmonic spectra of singlet and triplet H₂

- ightharpoonup Spectra at B=0 and at $B=B_0$ with and without (un)screened Lorentz contributions
 - restricted Hartree–Fock calculations in the uncontracted aug-cc-pVTZ London basis

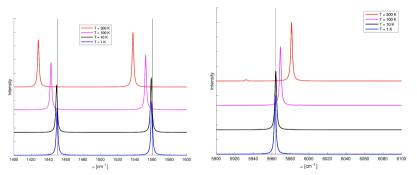


Lorentz force only affects modes that have a component perpendicular to the field vector:



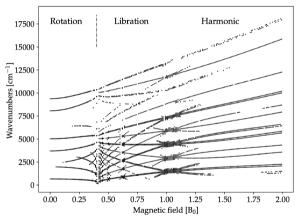
Harmonic and anharmonic frequencies of singlet H₂

- As the temperature of molecular dynamics calculations is lowered, the harmonic frequencies are recovered
 - this comparison represents an excellent test on our calculations



- Note: anharmonic corrections may be positive as well as negative
 - rotational mode becomes redshifted as rotation becomes less hindered at higher temperatures
 - stretching mode becomes blueshifted as more non-parallel orientations with a stiffer bond are explored at higher temperatures

Rovibrational spectrum of H₂ from Coulomb to Landau regime



► Three regimes: rotation (Coulomb), libration (intermediate), harmonic (Landau)

Conclusions and acknowledgements

Concluding remarks

- molecular dynamics in a magnetic field requires evaluation of Lorentz force
- screening of the Lorentz force calculated from the Berry curvature
- London atomic orbitals essential for accurate calculation of Berry curvature
- b diatomic spectra reveal hindered rotation and libration in the field
- Lorentz force responsible for splitting of otherwise degenerate levels
- harmonic frequencies and normal coordinates obtained from quadratic eigenvalue problem

Co-workers:

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- Wim Klopper, Ansgar Pausch, Christof Holzer, Laurenz Monzel (KIT)







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